### **NST** Instrumentation and Data







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New Solar Telescope



The New Solar Telescope

**Optical Configuration** 

The NST configuration is a 1/5 scale copy of one segment of the Giant Magellan Telescope.

The NST primary was figured by Steward Obs. Mirror Lab as a technology test for the GMT, greatly reducing mirror figuring costs for the NST project.

The NST secondary mirror was manufactured by Space Optics Research Labs.



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**Optical Configuration** 

The NST is an off-axis section of a 5.3 meter, F/0.73 parabolic parent with an elliptical secondary. The off-axis distance is 1.84 m resulting in a 1.6 meter, F/2.4 primary.

**Optical Parameters:** 

F/52 system EFL = 8.32 m Plate Scale = 2.48 arc sec/mm 180 arc sec field Gregorian image = 72.6 mm



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# **Ongoing Instrumentation**



- Adaptive Optics (AO-308)
- Visible Imaging Spectrometer (VIS)
- Cryogenic Infrared Spectrograph (CYRA)
- Near InfraRed Imaging
   Spectropolarimeter (NIRIS)





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Typical Raw Image taken with TiO filter at 7057A





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AO corrected Image taken with **TiO** filter at 7057A





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AO corrected + speckle Reconstructed Image taken with TiO filter at 7057A





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### **Visible Imaging Spectrometer**

- Single Fabry-Pérot etalon plus narrow band interference filter
- Spectrometer before NST AO-308 installation
- ✤ Wavelength coverage: 550 700 nm
- Band pass: 5.8 pm
- Telecentric optical configuration
- ✤ Field of view: 45" by 45"
- Available spectral lines:
  - Hlpha (656.3  $\pm$  0.15 nm
  - ✤ Fel(630.2 ± 0.15 nm)
  - $\bigstar \ \ \mathrm{NaD_2}$  (588.9  $\pm$  0.15 nm
- Spectrometry cadence:

✤ < 15 s</p>



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### VIS: 1<sup>st</sup> Light Observations







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### He I 1083 nm Observations





#### A microflare on May 24, 2012

A C9.5 flare on July 5, 2012



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### 2<sup>nd</sup> Generation -- NIRIS

- Dual Fabry-Pérot etalons with apertures of 100 mm
- Wavelength coverage: 1000– 1600 nm (He I 1083 and Fe I 1565 nm)
- New HgCdTe 2K by 2K camera with a frame rate up to 65 Hz
- Achromatic rotating waveplate
- Spectral resolving power:  $> 10^5$
- Telecentric optical configuration
- Field of view: 85"
- Polarimetry sensitivity:  $10^{-4} I_c$
- Spectroscopy cadence: < 5 s</p>
- Vector spectro-polarimetry cadence: < 10 s</li>



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# Evolution of Stokes







# Cryogenic IR Spectrograph





- Operating Temperature of 77 K to minimize background thermal emission
- Spectral range 1.0 to 5.0 μm
- ✤ Spectral resolving power > 10<sup>5</sup>
- Spatial resolution: diffraction-limited for wavelengths > 2 μm
- Temporal modulation 10 Hz with dual beam differential polarimetry
- Field of view: 75"
- Image stabilization: Tip/tilt
- Image rotation compensator
- image scanner, context imaging and guiding
- HAWAII 2RG 2K × 2K HgCdTe camera with a frame rate up to 65 Hz

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### Data Format



- Data are taken in a form of bursts. The burst size is 25 70 images and they are saved as one fits file
- One burst is acquired within 2 4 sec
- Each burst is later speckle reconstructed to produce one image
- Min cadence of the photospheric data is 12 sec; VIS (Halpha) 2 sec and more depending on the wavelength selection
- The data archive is not available online. Only a summary web page is generated and available on line
- The data can be requested using a web form. Zipped tar-balls will be available for download withing minuts to hours, depending on the data size



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### Please Enter the Date of the requested data





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#### Please fill in the form below



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#### The NST Data Form has been submitted!

This page will automatically check for the data every 10 seconds, please be patient as the process will take up to 30 min. If you do no wish to keep this tab open, simply wait for the email to arrive, it will also contain a ftp link to the data. If you have an ftp client setup, the link will open in that client, if not it will open in a new tab in your browser.

Checking again in: 00:05 seconds

### Waiting for data from TiO Cam. U

#### Waiting for data from H-alpha Cam. U



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### The NST Data Form has been submitted!

This page will automatically check for the data every 10 seconds, please be patient as the process will take up to 30 min. If you do no wish to keep this tab open, simply wait for the email to arrive, it will also contain a ftp link to the data. If you have an ftp client setup, the link will open in that client, if not it will open in a new tab in your browser.

#### **TiO Cam Data are Ready**

Click Here for FTP link to TiO Data

#### H-alpha Cam Data are Ready

Click Here for FTP link to H-alpha Data



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# Data Availability and Policy

- All TiO (photospheric) data are available upon request
- All VIS (H-alpha) data 9 month and older are available upon request
- Most recent (<9 mon) data: only H-alpha lie center, no off band
- Most recent (<9 mon) offband data are available only via PI's permission</li>
- CYRA and NIRIS data are not available thru the request form



### **Observations and Planning**



- May 1 July 31 Scheduled according to requests, both internal and external
- Aug 1 Sept 10 NAOC Observations
- Sept 15 April 30 patrol observations; no schedule
- Observe IRIS target whenever possible
- NST-IRIS Event list

